

## Chapter 43

6. We note that the sum of superscripts (mass numbers  $A$ ) must balance, as well as the sum of  $Z$  values (where reference to Appendix F or G is helpful). A neutron has  $Z = 0$  and  $A = 1$ . Uranium has  $Z = 92$ .

(a) Since xenon has  $Z = 54$ , then “Y” must have  $Z = 92 - 54 = 38$ , which indicates the element strontium. The mass number of “Y” is  $235 + 1 - 140 - 1 = 95$ , so “Y” is  $^{95}\text{Sr}$ .

(b) Iodine has  $Z = 53$ , so “Y” has  $Z = 92 - 53 = 39$ , corresponding to the element yttrium (the symbol for which, coincidentally, is Y). Since  $235 + 1 - 139 - 2 = 95$ , then the unknown isotope is  $^{95}\text{Y}$ .

(c) The atomic number of zirconium is  $Z = 40$ . Thus,  $92 - 40 - 2 = 52$ , which means that “X” has  $Z = 52$  (tellurium). The mass number of “X” is  $235 + 1 - 100 - 2 = 134$ , so we obtain  $^{134}\text{Te}$ .

(d) Examining the mass numbers, we find  $b = 235 + 1 - 141 - 92 = 3$ .

7. If  $R$  is the fission rate, then the power output is  $P = RQ$ , where  $Q$  is the energy released in each fission event. Hence,

$$R = P/Q = (1.0 \text{ W}) / (200 \times 10^6 \text{ eV})(1.60 \times 10^{-19} \text{ J/eV}) = 3.1 \times 10^{10} \text{ fissions/s.}$$

30. We are given the energy release per fusion ( $Q = 3.27 \text{ MeV} = 5.24 \times 10^{-13} \text{ J}$ ) and that a pair of deuterium atoms is consumed in each fusion event. To find how many pairs of deuterium atoms are in the sample, we adapt Eq. 42-21:

$$N_{d\text{pairs}} = \frac{M_{\text{sam}}}{2M_d} N_A = \left( \frac{1000 \text{ g}}{2(2.0 \text{ g/mol})} \right) (6.02 \times 10^{23} / \text{mol}) = 1.5 \times 10^{26}.$$

Multiplying this by  $Q$  gives the total energy released:  $7.9 \times 10^{13} \text{ J}$ . Keeping in mind that a watt is a joule per second, we have

$$t = \frac{7.9 \times 10^{13} \text{ J}}{100 \text{ W}} = 7.9 \times 10^{11} \text{ s} = 2.5 \times 10^4 \text{ y.}$$

36. The energy released is

$$\begin{aligned}
 Q &= -\Delta mc^2 = -(m_{\text{He}} - m_{\text{H}_2} - m_{\text{H}_1})c^2 \\
 &= -(3.016029 \text{ u} - 2.014102 \text{ u} - 1.007825 \text{ u})(931.5 \text{ MeV/u}) \\
 &= 5.49 \text{ MeV}.
 \end{aligned}$$

37. (a) Let  $M$  be the mass of the Sun at time  $t$  and  $E$  be the energy radiated to that time. Then, the power output is

$$P = dE/dt = (dM/dt)c^2,$$

where  $E = Mc^2$  is used. At the present time,

$$\frac{dM}{dt} = \frac{P}{c^2} = \frac{3.9 \times 10^{26} \text{ W}}{(2.998 \times 10^8 \text{ m/s})^2} = 4.3 \times 10^9 \text{ kg/s}.$$

(b) We assume the rate of mass loss remained constant. Then, the total mass loss is

$$\begin{aligned}
 \Delta M &= (dM/dt) \Delta t = (4.33 \times 10^9 \text{ kg/s}) (4.5 \times 10^9 \text{ y}) (3.156 \times 10^7 \text{ s/y}) \\
 &= 6.15 \times 10^{26} \text{ kg}.
 \end{aligned}$$

The fraction lost is

$$\frac{\Delta M}{M + \Delta M} = \frac{6.15 \times 10^{26} \text{ kg}}{2.0 \times 10^{30} \text{ kg} + 6.15 \times 10^{26} \text{ kg}} = 3.1 \times 10^{-4}.$$

51. Since plutonium has  $Z = 94$  and uranium has  $Z = 92$ , we see that (to conserve charge) two electrons must be emitted so that the nucleus can gain a  $+2e$  charge. In the beta decay processes described in Chapter 42, electrons and neutrinos are emitted. The reaction series is as follows:

